

PLASMA LOSSES

Plasma Internal Energy Losses

As most of the fuel is lost unreacted

Plasma Radiation Losses

Radiative Recombination

Bremsstrahlung Radiation

when charged particles are decelerated by collision with other charged particles and emit a photon

Dielectronic Radiation

Spectral line Radiation

Atomic Line Radiation

result from a 3-body capture of a free electron by an ion with a third body in the vicinity. Binding energy of the electron is radiated.

results if an impurity is raised to an excited state by electron impact.

Cyclotron Radiation

emitted from electrons when they gyrate in a strong magnetic field

The capture of a free electron with simultaneous excitation of a second orbital electron, such that both electrons are in an excited state.

Plasma Internal Energy Losses

in a fusion reactor, only $\sim 20\%$ of the fuel is burned-up. Thus, most of the fuel is lost unreacted. The energies carried by the unburned fuel by charged reaction products and the external heating of the fuel.

The rate of internal energy loss :

$$P_E = \frac{3}{2} \sum_j \frac{n_j K T_j}{\tau_j}$$

$\parallel \tau_j$ = particle containment time

$$= \frac{3}{2} n_1 \frac{K T_i}{\tau_i} + \frac{3}{2} n_2 \frac{K T_e}{\tau_e} + \frac{3}{2} n_{Pr} \frac{K T_{Pr}}{\tau_{Pr}} + \frac{3}{2} n_e \frac{K T_e}{\tau_e}$$

un-reacted fuel
confined in the
plasma

charged
reaction
Products
confined
in the plasma

electron
population

Consider: $\tau_i, \tau_e, \tau_{Pr}$ are long, so
ionic species equilibrate to the
same kinetic temperature

$$= \frac{3}{2} \frac{K T_i}{\tau_i} (n_1 + n_2 + n_{Pr}) + \frac{3}{2} \frac{K T_e}{\tau_e} n_e$$

For steady state operation "Quasi-neutral":

$$\tau_i \approx \tau_e \approx \tau_{pr}$$

if the burn-up fraction is small,

$$\text{i.e. } n_{pr} \ll n_1 \\ n_{pr} \ll n_2$$

Then $n_i \approx n_1 + n_2$

if containment time is long enough for ions and electrons to equilibrate, i.e. $\tau_i \approx \tau_e$

so: $P_E = \frac{3}{2} \frac{KT}{\tau_{pr}} (n_1 + n_2 + n_e)$

In a fully-ionized plasma with two fuel species:

$$n_e = Z_1 n_1 + Z_2 n_2 \quad \left| \begin{array}{l} Z_1 \text{ and } Z_2 \text{ are} \\ \text{the proton \# of} \\ \text{fuel species} \end{array} \right.$$

$\frac{n_1}{n_i} + \frac{n_2}{n_i} = 1$

$$\frac{n_1}{n_i} = \gamma, \frac{n_2}{n_i} = 1 - \gamma$$

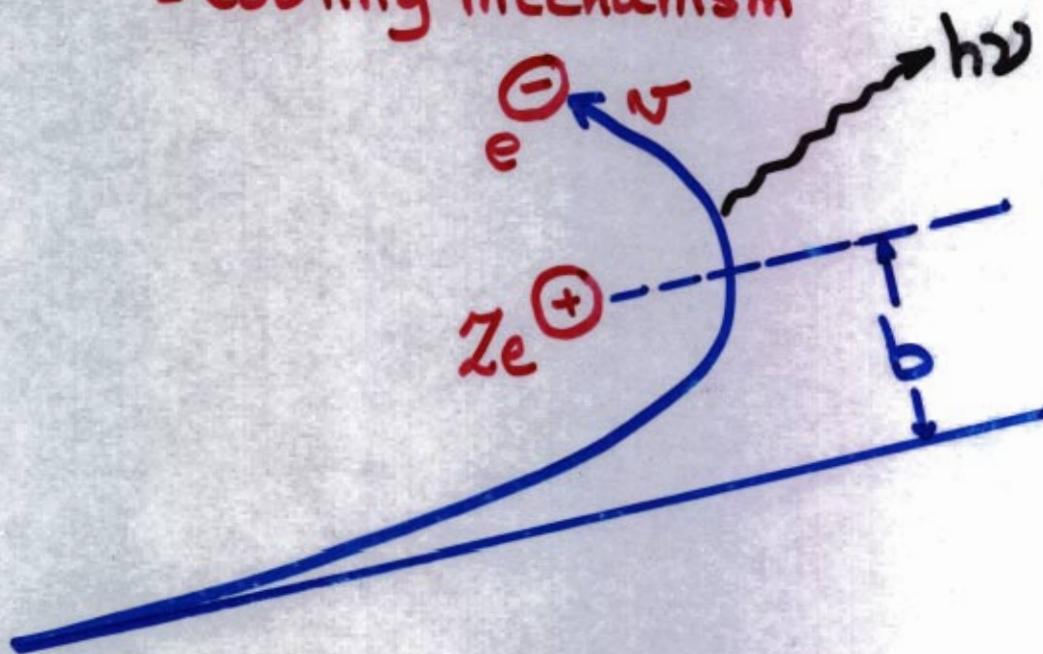
$$\left| \begin{array}{l} \frac{n_1}{n_i} = \gamma = \text{fuel mixture} \\ \text{ratio} \\ \equiv \frac{\text{1st fuel species}}{\text{Total fuel}} \end{array} \right.$$

hence:

$$P_E = \frac{3}{2} \frac{KT}{\tau_{pr}} [1 + \gamma Z_1 + (1 - \gamma) Z_2] n_i \text{ w/m}^3$$

Bremsstrahlung Radiation

Causes :- Energy loss
- Cooling mechanism



from Classical Electromagnetic theory :-
U; Total power radiated from a single electron
of charge e, velocity v, acceleration \ddot{v}

$$\frac{dU}{dt} = \frac{e^2}{6\pi\epsilon_0 c^3} \frac{[|\dot{v}|^2 - (v \times \dot{r})^2/c^2]}{(1-v^2/c^2)^3} \text{ watts}$$

non-relativistic limit : $v \ll c$

$$\frac{dU}{dt} = \frac{e^2 \dot{v}^2}{6\pi\epsilon_0 c^3} \text{ watts}$$

The Coulomb force $F_C = \frac{Ze^2}{4\pi\epsilon_0 b^2} = m_e \dot{v}$

$$\text{hence: } \dot{U} = \frac{Z e^2}{4\pi \epsilon_0 m_e b^2}$$

Consequently :

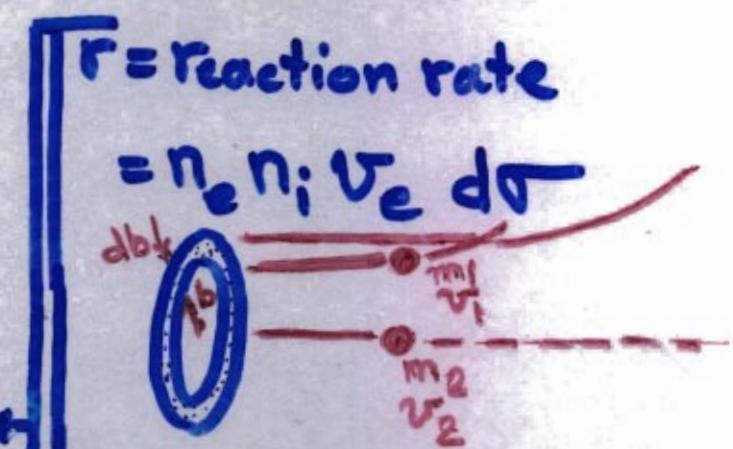
$$\frac{dU}{dt} = -\frac{Z^2 e^6}{96 \pi^3 \epsilon_0^3 C^3 m_e^2 b^4}$$

and $b = t_0 v_e \parallel t_0$: electron is in the vicinity of the ion for time t_0 .
The energy radiated / collision:

$$E_R \approx t_0 \frac{dU}{dt} = \frac{b}{v_e} \frac{dU}{dt} \text{ Joules}$$

TOTAL Radiated Power:

$$\begin{aligned} dP_{Br} &= E_R \times \Gamma \\ &= 2\pi n_e n_i \frac{dU}{dt} b^2 db \\ &= \frac{Z^2 e^6 n_e n_i}{48 \pi^2 \epsilon_0^3 C^3 m_e^2} \frac{db}{b^2} \end{aligned}$$

$\Gamma = \text{Reaction rate}$
 $= n_e n_i v_e d\sigma$

 $d\sigma = 2\pi b db$

$$P_{Br} = \frac{Z^2 e^6 n_e n_i}{48 \pi^2 \epsilon_0^3 C^3 m_e^2} \int_{b_{min}}^{b_{max} \rightarrow \infty} \frac{db}{b^2} w/m^3$$

$$= \frac{Z^2 e^6 n_e n_i}{48 \pi^2 \epsilon_0^3 C^3 m_e^2} b_{min} w/m^3$$

replace b_{min} by the Compton Wavelength for electrons : $b_{min} = \lambda_c = \frac{h}{2\pi m_e v_e}$

$$P_{Br} = \frac{n_e n_i Z^2 e^6 v_e}{24 \pi \epsilon_0^3 C^3 m_e h}$$

for Maxwellian plasma

$$f(v) = \frac{dn}{dv}$$

$$= \frac{4n}{\pi^{1/2}} \left(\frac{m}{2kT} \right)^{3/2} v^2 \exp \left(-\frac{mv^2}{2kT} \right)$$

The mean thermal velocity :

$$\text{"electrons"}: \bar{v} = \frac{1}{n} \int_0^\infty v f(v) dv = \left(\frac{8kT_e}{\pi m_e} \right)^{1/2}$$

cannot replace by λ_D as per the case for the rate of change of momentum, it is NOT a shielding case as a photon is emitted

$$P_{Br} = \frac{n_e n_i Z^2 e^6}{24\pi \epsilon_0^3 C^3 m_e h} \left(\frac{8KT_e}{\pi m_e} \right)^{1/2} \text{ w/m}^3$$

From Quantum mechanical calculations,
the result is larger by a factor $(\frac{3}{2})^{1/2}$

$$\begin{aligned} P_{Br} \Big|_{Q.M.} &= \frac{e^6}{24\pi \epsilon_0^3 C^3 m_e h} \left(\frac{12KT_e}{\pi m_e} \right)^{1/2} n_e n_i Z^2 \\ &= 1.625 \times 10^{-38} n_e n_i Z^2 \frac{T_e^{1/2}}{(\text{eV})} \end{aligned}$$

* for two fuel species:

$$\text{replace } n_i Z^2 \text{ by } \sum_j n_j Z_j^2$$

* for quasi-neutral plasmas:

$$\text{replace } n_e \text{ by } \sum_j n_j Z_j$$

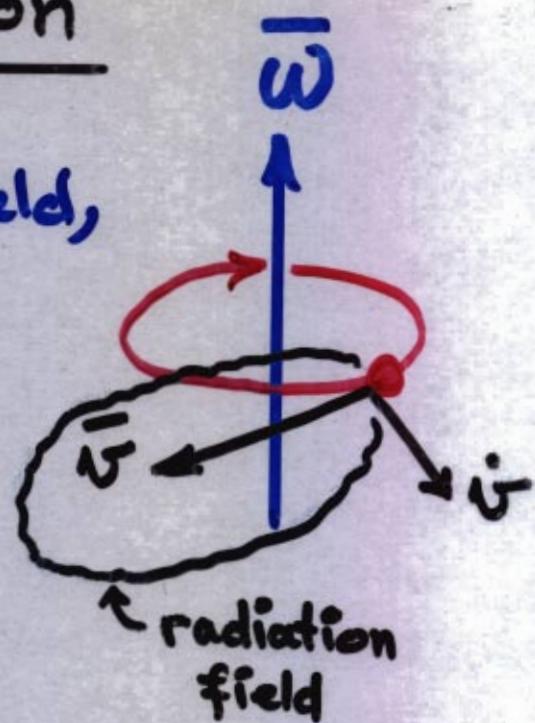
$$P_{Br} = 1.625 \times 10^{-38} \frac{T_e^{1/2}}{(\text{eV})} \sum_j n_j^2 Z_j^3$$

for Advanced and Aneutronic fusion reaction;
high Z-isotopes, which is a restriction.

Bremsstrahlung is a significant loss for DT
and more serious for higher Z fuels.

Cyclotron Radiation

in a strong magnetic field, an electron will gyrate with frequency ω . If electron is energetic, the acceleration will induce the emission of photons from the electron in the direction of its velocity \vec{v} .



As per classical Electromagnetic theory:

$$\frac{dU}{dt} = \frac{e^2}{6\pi\epsilon_0 c^3} \frac{[|\vec{v}|^2 - (\vec{v} \times \vec{\omega})^2/c^2]}{(1 - v^2/c^2)^3} \text{ watts}$$

The relativistic cyclotron frequency is:

$$\omega = \frac{eB}{m_e} \left(1 - \frac{v^2}{c^2}\right)^{1/2}$$

and the acceleration is: $\vec{U} = \omega \vec{v}$

Since only \dot{U}_\perp contributes, as $\dot{U} \perp \omega$

hence:

$$[|\dot{U}|^2 - (\dot{U} \times \dot{U})^2/c^2] =$$

$$= \omega^2 U_\perp^2 - \frac{\omega^2 U_\perp^4}{c^2} = \omega^2 U_\perp^2 \left(1 - \frac{U_\perp^2}{c^2}\right)$$

Substitute into dU/dt :

$$\frac{dU}{dt} = \frac{e^4 U_\perp^2 B^2}{6\pi\epsilon_0 c^3 m_e^2 (1 - U_\perp^2/c^2)} \text{ watts}$$

[for a single electron]

$$\left. \frac{dU}{dt} \right|_{\substack{\text{all} \\ \text{electrons}}} = n_e \frac{dU}{dt}$$

Consequently; for a Maxwellian plasma

$$P_{\text{cyc}} = \frac{e^4 B^2 n_e}{3\pi\epsilon_0 m_e^2 c} \left(\frac{K T_e}{m_e c^2} \right) \left[1 + \frac{5 K T_e}{2 m_e c^2} + \dots \right]$$

w/m³

*N.B. integration is carried-out for an optically thin plasma

$$P_{\text{cyc}} = 6.21 \times 10^{-20} n_e T_e \frac{B^2}{(\text{ev})} \left(1 + \frac{T_e (\text{ev})}{2.04 \times 10^5} + \dots \right) \text{W/m}^3$$

The emitted photons from a fusion plasma have frequencies $\sim 100 \text{ GHz}$, in the far infrared, and absorbed in the plasma. also, they are reflected from the metallic walls.

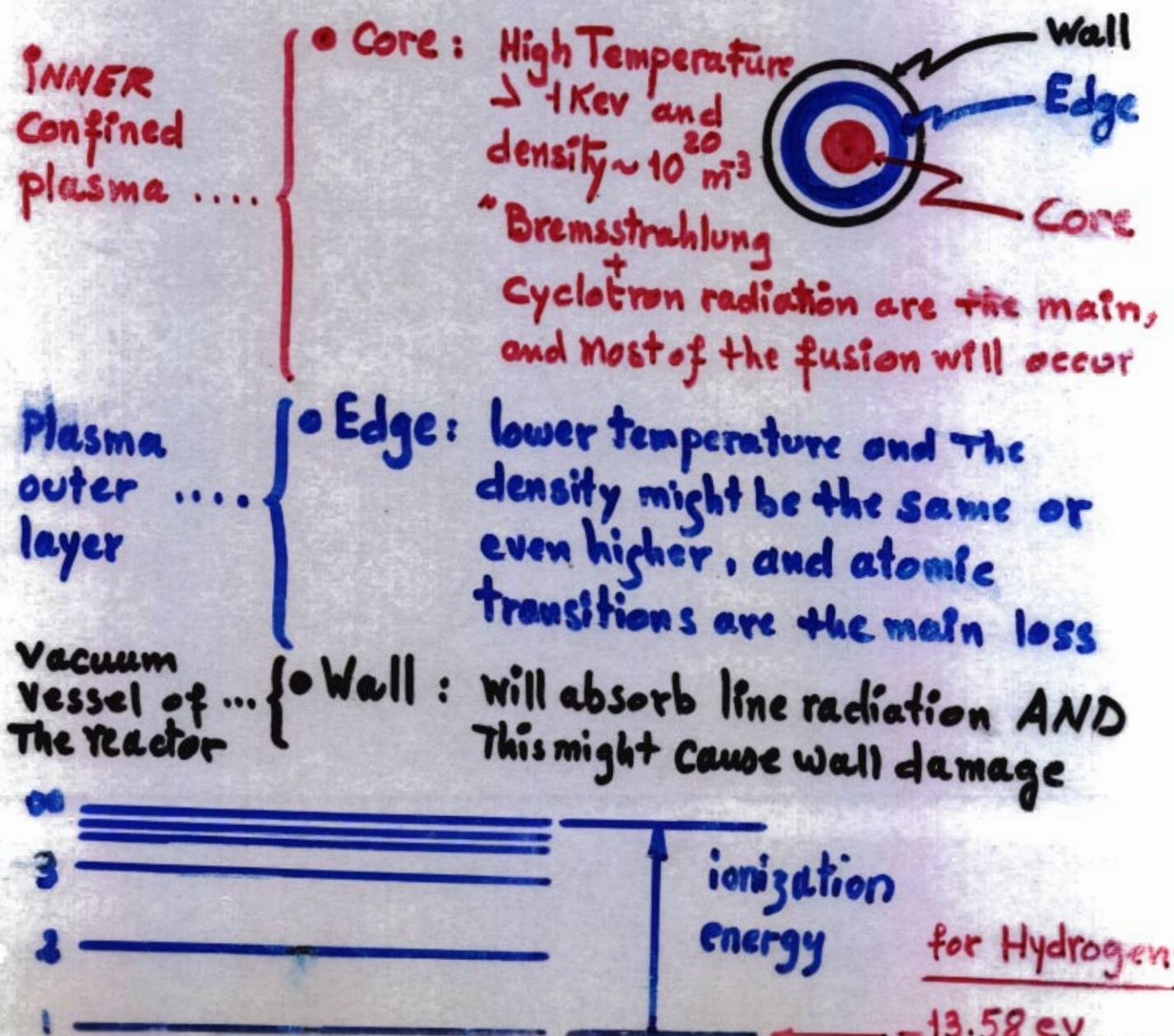
Radiation transport calculations showed that the plasma core radiates strongly So the plasma will be heated near the outer radius.

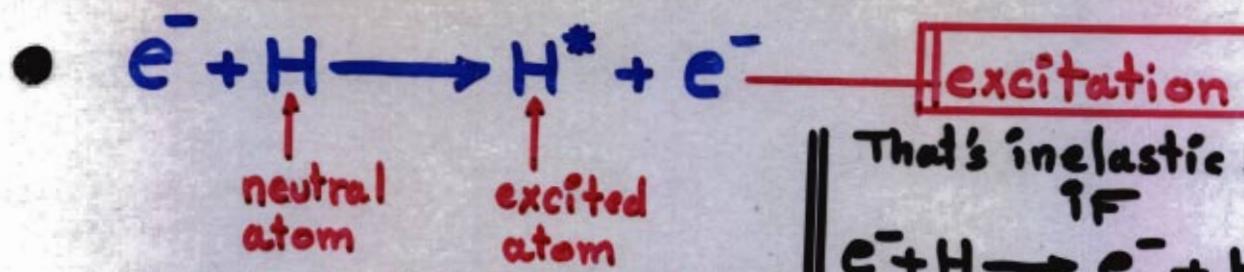
Thus: Causes:

- * Cooling of the Core
- * heating of outer periphery

Charge exchange, Ionization, Impurities and Line Radiation

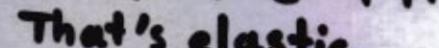
This is directly related to the Atomic Collision Phenomena. Atomic and Molecular Collisions are responsible for the charge exchange which will cause the loss of hot ions. The existence of impurities will absorb the electron energy through ionization, and consequently will emit line radiation which will be mainly absorbed by the fusion reactor walls.



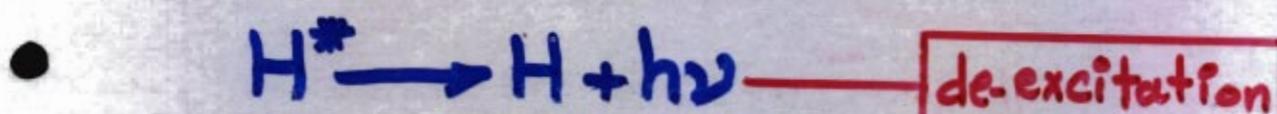


Takes place when the electron populates one of the excited states, i.e. the atom will be overpopulated by an extra electron.

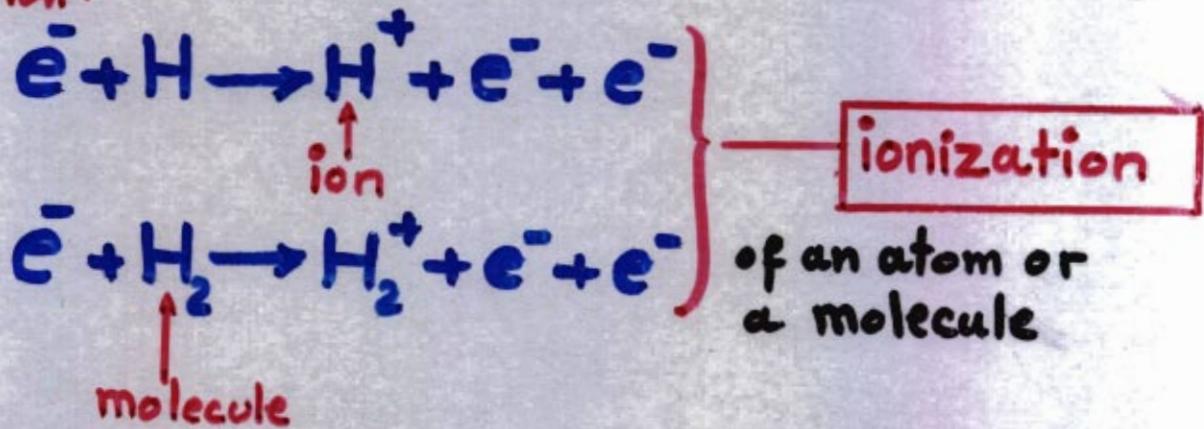
That's inelastic scattering
IF $e^- + H \longrightarrow e^- + H$



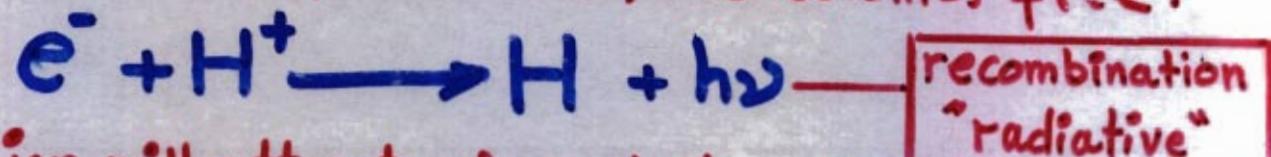
That's elastic



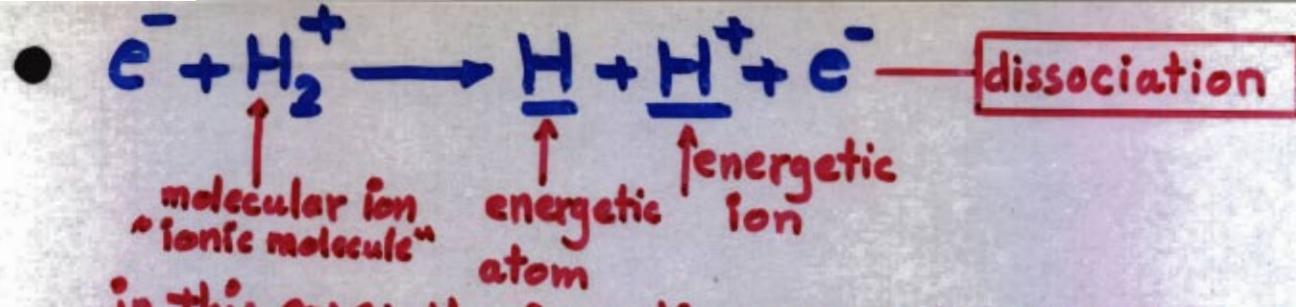
The excited atom is de-excited to the ground state and the ionization energy is released in the form of a photon.



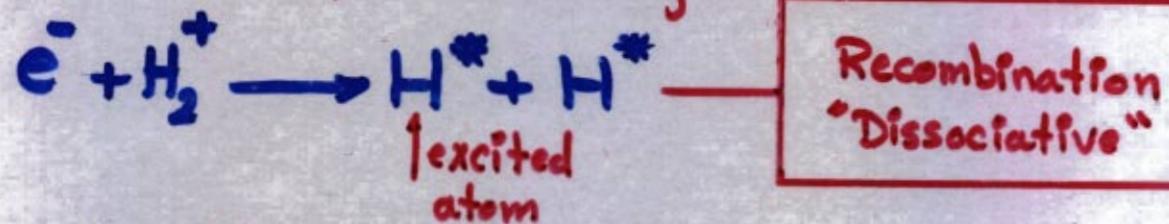
This takes place when the electron transfers enough energy to the bound electron, so as the latter will leave the atom (or molecule) and becomes free.



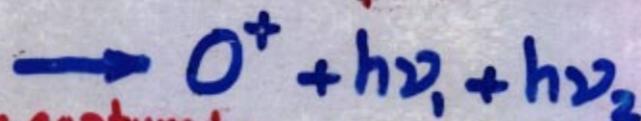
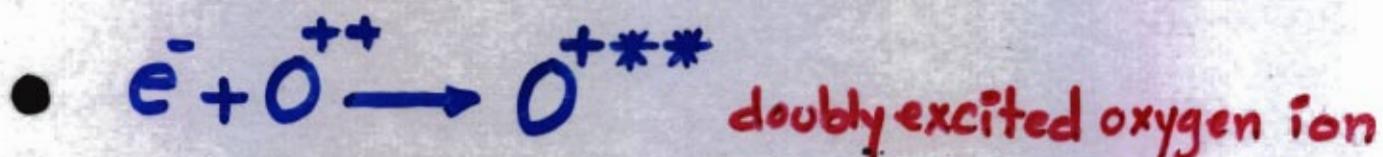
The ion will attract a free electron to the ground state and RECOMBINE to form a neutral atom, and the binding energy will be released in the form of a photon "radiation".



In this case; the formation of an energetic atom and an energetic ion will take place, which means that an atom will be formed, and if:-

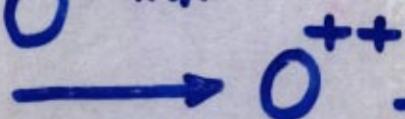
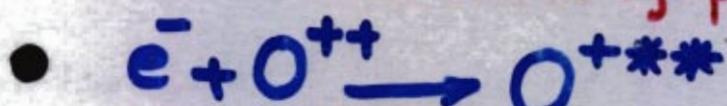


excited atoms are formed, the dissociation is called DISSOCIATED RECOMBINATION



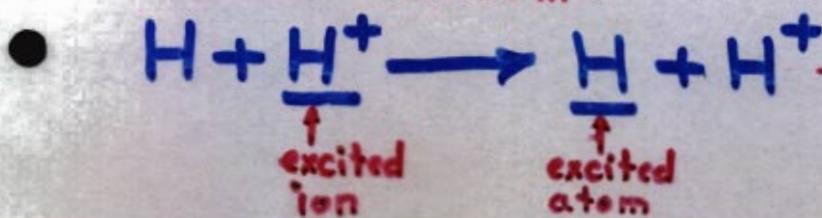
The free electron is captured +
 Simultaneous excitation, which
 results in the formation of an
 ion and the release of photons.

Recombination
 "dielectronic"



This will take place when one of the
 excited electrons acquires enough energy
 to leave the atom.

autoionization



charge exchange

usually occurs for ions and
 atoms of like species.

EQUILIBRIUM

Consider the atom density in an excited state 1 to be n_1 , in state 2 to be n_2 ; and consider W_1 and W_2 to be the corresponding energy states. According to Maxwell-Boltzmann distribution for excited states:

$$\frac{n_1}{n_2} = \frac{g_1}{g_2} \exp\left\{\frac{-(\Delta W)}{kT}\right\}$$

g : is the degeneracy factor

{ For the case of }
 { high-density }
 { Plasmas }
 state 2: neutral atom
 state 1: electron-ion pair

$$\begin{aligned} \Delta W &= \\ W_1 - W_2 &= \\ &\text{ionization Potential} \end{aligned}$$

then: $\frac{g_1}{g_2} = \left(\frac{2\pi m_e k T}{h^2}\right)^{3/2} \cdot \frac{1}{n_1}$

hence:
 ion density $\leftarrow \frac{n_1^2}{n_2}$ $= \left(\frac{2\pi m_e k T}{h^2}\right)^{3/2} \exp\left(-\frac{\Delta W}{kT}\right)$
 neutral atom density $\leftarrow \frac{n_1}{n_2}$

hence:

$$\frac{n_i}{n_n} = 3.022 \times 10^{27} \frac{T_{ev}^{-3/2}}{n_i} \exp\left(-\frac{\Delta W}{T}\right)$$

SAHA equation

The equation tells the expected amount of ionization in a gas in thermal equilibrium

The fractional ionization is given by $\frac{n_i}{n_i + n_n}$
 $\approx \frac{n_i}{n_n}$

$$\parallel n_i \ll n_n$$

To understand the meaning

we can consider Air at room temperature of 300°K (0.026 eV) with $\Delta W = 14.5\text{ eV}$ for nitrogen, and n_n of the value $3 \times 10^{25}\text{ m}^{-3}$.

This gives a fractional ionization :

$$\frac{n_i}{n_n} \approx 10^{-122}$$

with temp. increase, the degree of ionization remains low until ΔW becomes few times T , then $\frac{n_i}{n_n}$ rises abruptly and the gas becomes in a plasma state. With further increase in T , n_n becomes $< n_i$ and the plasma becomes FULLY IONIZED

In high-density plasmas, Collisional processes may dominate de-excitation and recombination and the relative populations in Thermodynamic equilibrium are governed by the Saha eqn., depending only on Temperature.

if the total heavy particle density ($n_t = n_i + n_n$) is introduced to the Saha eqn., then:

$$\frac{n_i}{n_t} \approx 1 - \left[3.313 \times 10^{-28} n_t T_{\text{ev}}^{-3/2} e^{\left(\frac{\Delta W}{T} \right)} \right]$$

let us consider hydrogen gas at $T = 10 \text{ ev}$ at a pressure of $5 \times 10^3 \text{ atm}$.

$$n_t = \frac{P}{kT} \quad \parallel P \text{ in Pascal or Newton/m}^2 \\ \text{Joules} \quad \parallel 1 \text{ atm} = 1.013 \times 10^5 \text{ Pa} \\ \parallel 1 \text{ Torr} = 133.3 \text{ Pa}$$

$$P = 5 \times 10^3 \times 1.013 \times 10^5 = 506.5 \text{ Pa}$$

$$kT = 10 \times 11609 \times 1.38 \times 10^{-23} = 1.6 \times 10^{-18}$$

$$n_t = 3.166 \times 10^{20} \text{ m}^{-3}$$

$$\frac{n_i}{n_t} = 1 - \left[3.313 \times 10^{-28} \times 3.166 \times 10^{20} \times (10)^{-3/2} e^{\frac{13.6}{10}} \right] \\ = 1 - 1.29 \times 10^{-5} \approx 1 \text{ i.e. } n_i \approx n_t$$

100% ionization \Rightarrow Fully ionized

if $T = 1 \text{ ev}$; $n_t = 3.166 \times 10^{21} \text{ m}^{-3}$

and $\frac{n_i}{n_t} = 84.5\%$ ionization

if $T = 0.1 \text{ ev} \rightarrow \approx 0\%$ ionization ($1 - 3.8 \times 10^{52}$)

To the Solar System :

(hot plasma with low density)

there is a balance between :-

[rate of Collisional excitation]
[rate of ionization]

AND

[rate of radiative de-excitation]
[rate of radiative recombination]

Determines
the
populations
of
excited
levels

CORONAL EQUILIBRIUM

i.e. depends on the various rate of coeff. $\langle \sigma v \rangle$

For a hot plasma (i.e. high temp.), at the Core :

$$\text{rate of recombination} = n_e n_i \langle \sigma_F v_e \rangle$$

$$\text{rate of ionization by electrons} = n_e n_n \langle \sigma_e v_e \rangle$$

$$\text{rate of ionization by ions} = n_i n_n \langle \sigma_i v_i \rangle$$

[$n_n \rightarrow$ neutral density]

hence:

$$n_e n_i \langle \sigma_F v_e \rangle = n_e n_n \langle \sigma_e v_e \rangle +$$

$$+ n_i n_n \langle \sigma_i v_i \rangle$$

$$\frac{n_i}{n_n} = \frac{\langle \sigma_e v_e \rangle + \langle \sigma_i v_i \rangle}{\langle \sigma_F v_e \rangle}$$

at the Core:
neglect the rate of
Collisional excitation
as well as the rate
of radiative de-excit.

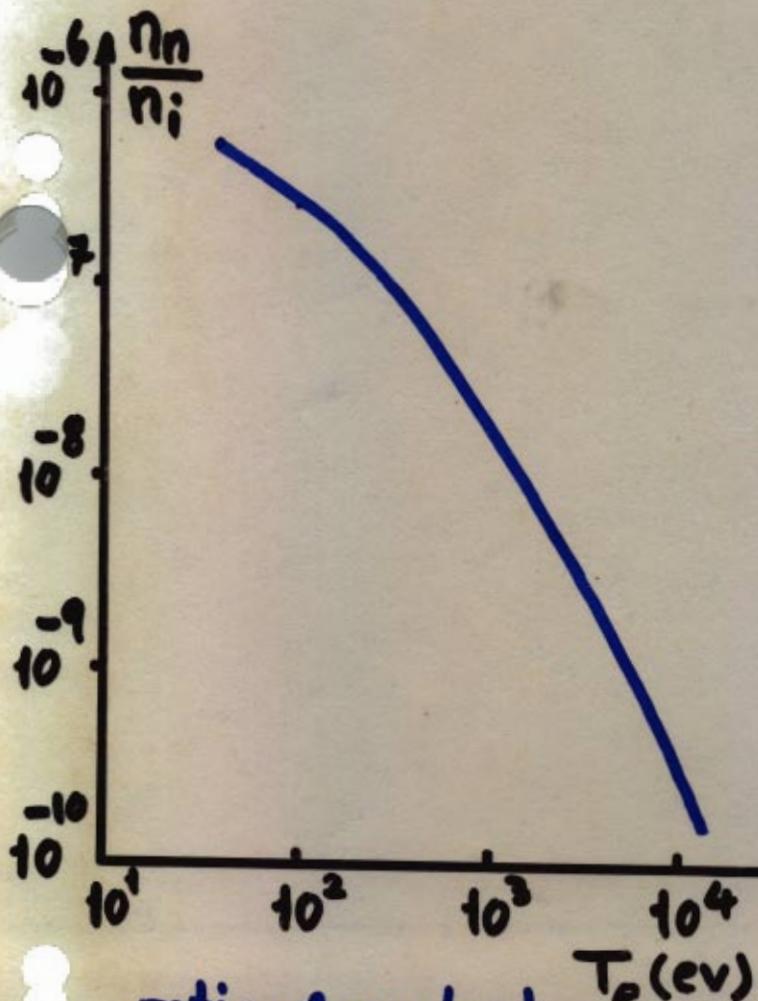
hot-low density
plasma : $n_e \approx n_i$

and the fractional Concentration = $\frac{n_n}{n_i + n_n}$

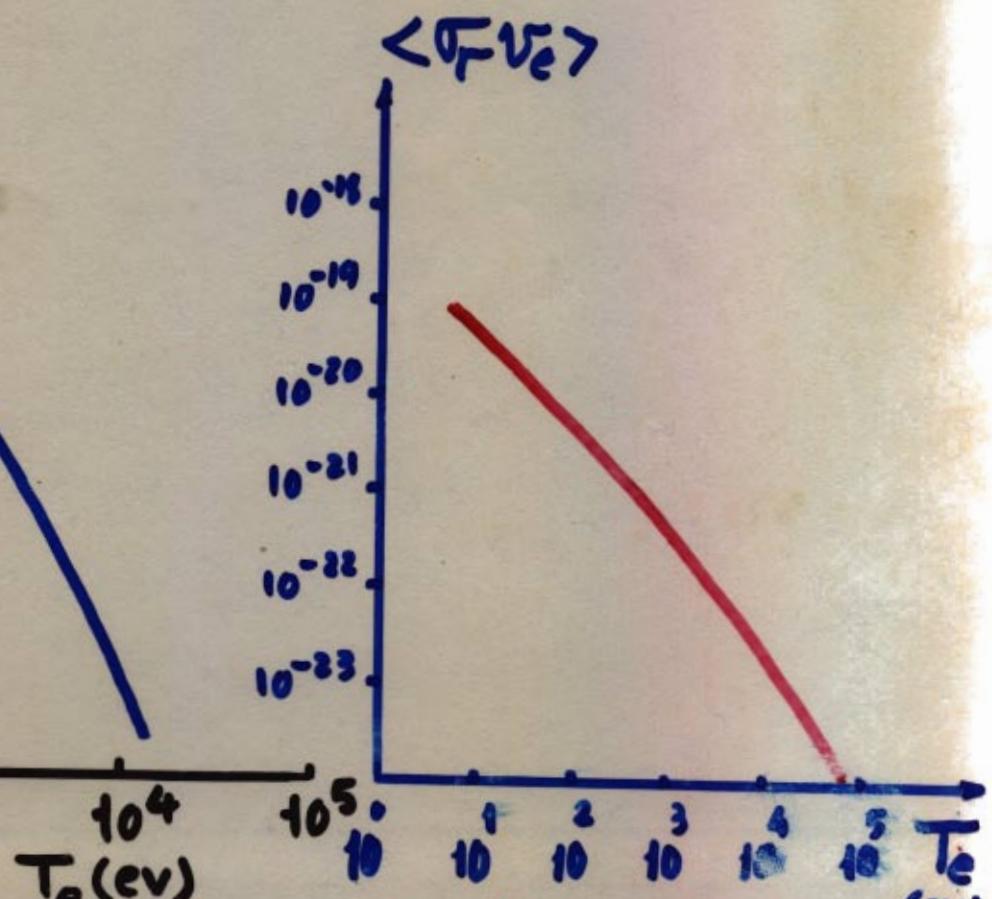
$$\frac{n_n}{n_i + n_n} = \frac{\langle \sigma_F v_e \rangle}{\langle \sigma_F v_e \rangle + \langle \sigma_e v_e \rangle + \langle \sigma_i v_i \rangle}$$

$$\approx \left[\frac{1}{1 + \frac{\langle \sigma_e v_e \rangle}{\langle \sigma_F v_e \rangle}} \right] \quad \begin{array}{l} \langle \sigma_i v_i \rangle \text{ very} \\ \text{small, could be} \\ \text{neglected} \end{array}$$

i.e. depends on the temp. as $\langle \sigma_e v_e \rangle$ and $\langle \sigma_F v_e \rangle$ are functions of temp.



ratio of neutral
atom density to
plasma density



radiative recombination
rate coefficient for
Hydrogen

I.2 PLASMA RADIATION LOSSES

Bremsstrahlung
Radiation

Deceleration of
Charged particles
by Collision
with Other
Charged particles



Cyclotron
Radiation

emission of
photons due
to electron
gyration in
strong
magnetic
fields

Atomic
Line
Radiation

- Radiative recombination
- Spectral line radiation
- Dielectronic radiation

Bremsstrahlung :

Energy Loss
Cooling Mechanism

$$P_{Br} = \left(\frac{e^6}{24\pi \epsilon_0^3 c^3 m_e h} \right) \left(\frac{12kT_e}{\pi m_e} \right)^{1/2} n_e n_i Z^2$$

Two fuel species: $n_i Z^2 \Rightarrow \sum_j n_j Z_j^2$

Quasi-neutral plasmas: $n_e \Rightarrow \sum_j n_j Z_j$

FINAL
↓

$$P_{Br} = 1.625 * 10^{-38} \frac{T_e^{1/2}}{(eV)} \sum_j n_j^2 Z_j^3$$

$$\cong 5.35 * 10^{-37} n_e n_i Z^2 T_e^{1/2}$$

$(m^{-3}) (m^{-3})$ (keV)

$$\frac{\Delta P_{Br}}{\text{ratio with impurities}} = 1 + \gamma (Z + Z^2) + \gamma^2 Z^3$$

\downarrow impurity fraction \downarrow impurity Z.

Cyclotron:

Core Cooling Outer periphery heating

$$P_{\text{cyc}} = \frac{e^4 B^2 n_e}{3\pi\epsilon_0 m_e^2 c} \left(\frac{kT_e}{m_e c^2} \right) \left(1 + \frac{5kT_e}{2m_e c^2} + \dots \right)$$

$$P_{\text{cyc}} \approx 6.21 \times 10^{-17} n_e T_e B^2 \left(1 + \frac{T_e \text{ (KeV)}}{146} \right)$$

(W/m³) (m⁻³) (KeV) (Tesla)

Atomic Line:

- Spectral line radiation :

$$P_{LR} \approx 8 \times 10^{-35} n_e n_i Z^6 T_e^{3/2}$$

(W/m³) (m⁻³)(m⁻³) (KeV)

Dominant factor at lower plasma temperature

- Radiative Recombination:

$$P_{RC} \approx 6 \times 10^{-40} n_e n_i Z^4 T_e^{1/2}$$

(W/m³) (m⁻³)(m⁻³) (KeV)

effective at higher temp.

The NON-radiative losses belong to the branch of INTERNAL ENERGY LOSSES due to the fact that only 20% of the fuel will be burned, the remaining un-burned fuel is then Lost unreacted.

$$P_E = 2.44 * 10^{-16} [1 + \gamma Z_1 + (1-\gamma) Z_2] \frac{n_i T}{\tau_{Pr}}$$

fuel ratio $n_i \Rightarrow m^{-3}$
 $T \Rightarrow \text{KeV}$
 $\tau \Rightarrow \text{sec} \approx \tau_i \approx \tau_e$
 Pr for Steady State operation

Low plasma temperature: P_{LR} is dominant

high plasma temperature: P_{RC} & P_{BR}

Very high plasma temperature: P_{BR}

Total Radiation Power Loss

$$\underbrace{P_{BR} + P_{LR} + P_{RC}}_{\text{function of } n_e \text{ and } n_{\text{impurities}} \downarrow (n_K)} = + K_{\text{cyc}} P_{\text{cyc}}$$

fraction absorbed in the walls

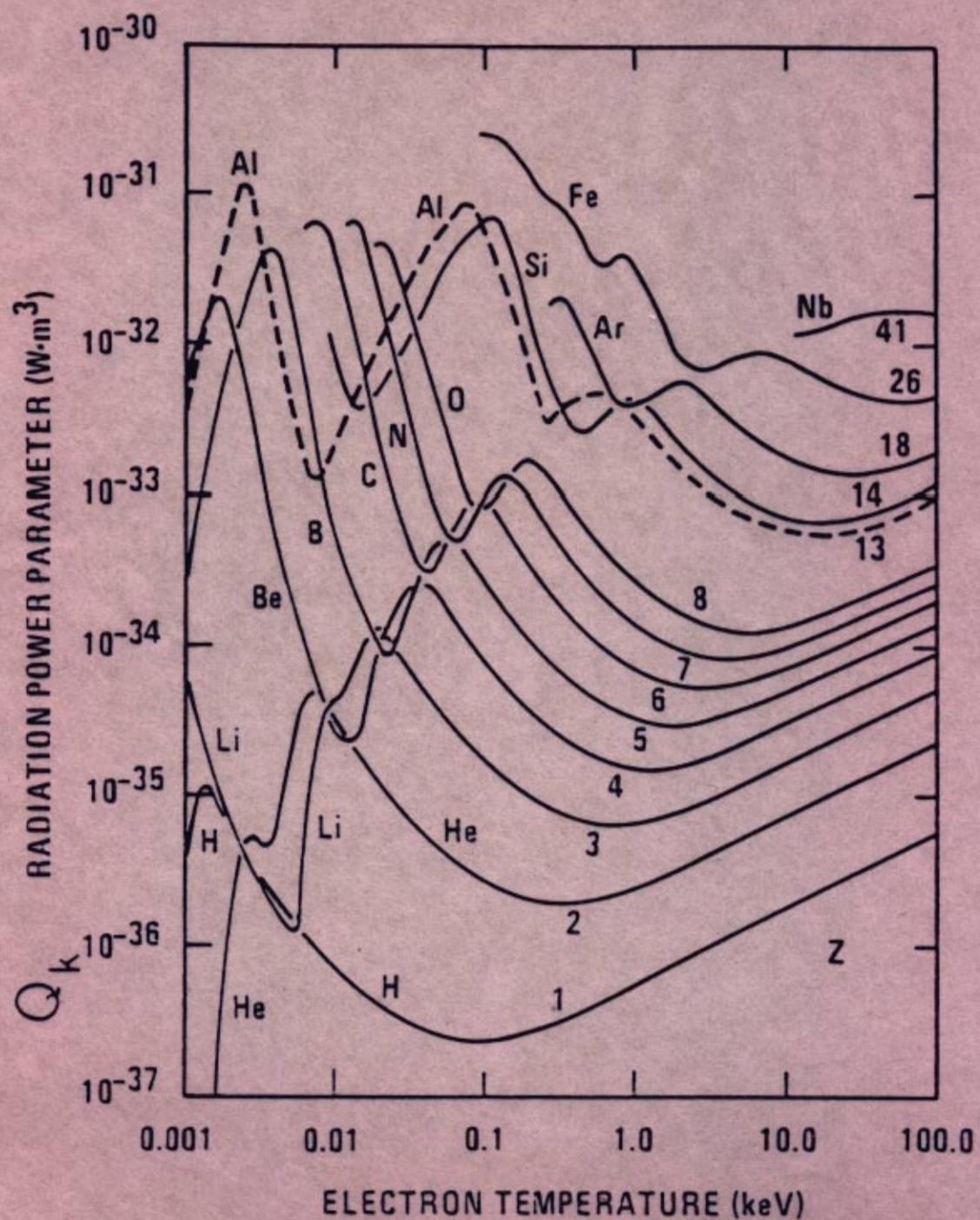
\therefore Radiation Power Parameter $Q_K =$

$$= \frac{P_{BR} + P_{LR} + P_{RC}}{n_e n_K}$$

i.e. $P_{RAD} = \sum_K n_e n_K Q_K + K_{\text{cyc}} P_{\text{cyc}}$

from tables and
Curves, or
Calculate and
SUM-UP

To be
Calculated
Separately



The radiation power parameter is a function of electron temperature for various elements. (Cyclotron radiation losses must be computed separately.) From J. R. Hopkins and J. M. Bowles, Nuclear Technology 43, 532 (1973), Fig. 1.

To calculate K_{cyc} :

$$K_{cyc} = f(T_e, D)$$



↑ plasma depth

Available in tables
and Curves for
Cylindrical plasmas

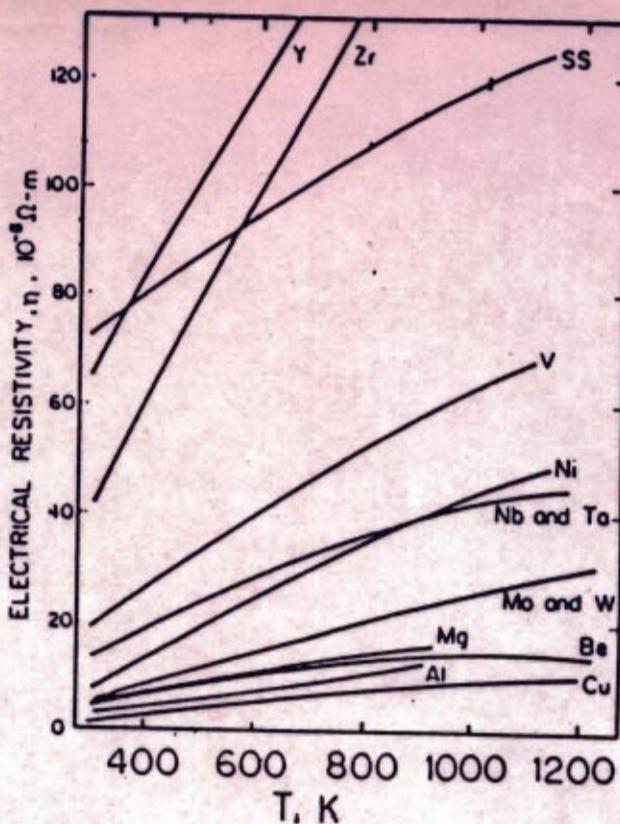
$$D = \frac{2\pi e a}{B^{3/2} \eta^{1/2} (1-\beta)^{3/4}}$$

a = plasma radius (m)

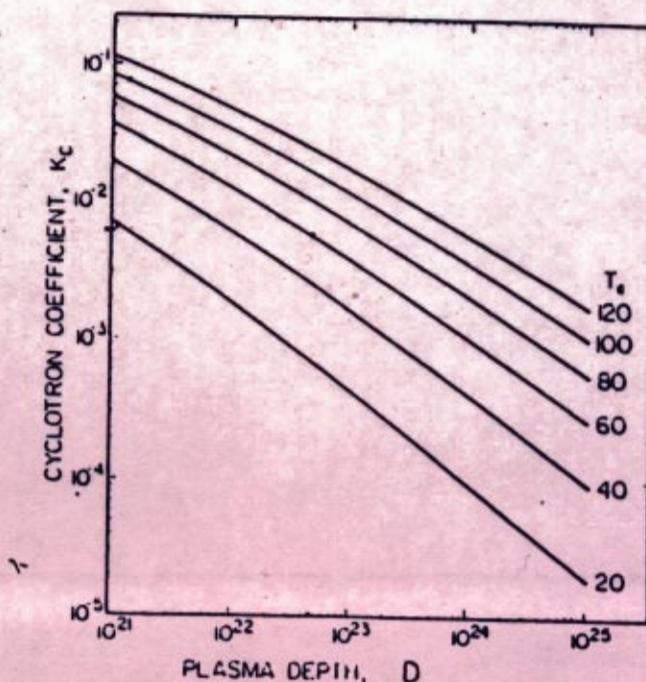
η = material resistivity ($\Omega \cdot m$)

$$\beta = \frac{P}{B^2/2\mu_0} = \frac{\text{plasma Kinetic Pressure}}{\text{magnetic pressure}}$$

- Steps:
- get η at the Wall temperature
 - get β from $P = \sum_j \eta_j k T_j$ & B
 - get D
 - use the curves.



Electrical resistivity vs. temperature for various metals. From G. H. Miley, *Fusion Energy Conversion*, American Nuclear Society, LaGrange Park, IL, 1975.



Cyclotron radiation coefficient K_C vs. "plasma depth" D , for a cylindrical plasma with various T_e (keV).